# Markscheme 

May 2016

Physics

Higher level

## Paper 3

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## Subject Details: Physics HL Paper 3 Markscheme

## Mark Allocation

## Candidates are required to answer ALL questions in Section A [15 marks] and all questions from ONE option in Section B [30 marks]. Maximum total = [45 marks].

1. Each row in the "Question" column relates to the smallest subpart of the question.
2. The maximum mark for each question subpart is indicated in the "Total" column.
3. Each marking point in the "Answers" column is shown by means of a tick $(\checkmark)$ at the end of the marking point.
4. A question subpart may have more marking points than the total allows. This will be indicated by "max" written after the mark in the "Total" column. The related rubric, if necessary, will be outlined in the "Notes" column.
5. An alternative wording is indicated in the "Answers" column by a slash ( $/$ ). Either wording can be accepted.
6. An alternative answer is indicated in the "Answers" column by "OR" between the alternatives. Either answer can be accepted.
7. Words in angled brackets «" in the "Answers" column are not necessary to gain the mark.
8. Words that are underlined are essential for the mark.
9. The order of marking points does not have to be as in the "Answers" column, unless stated otherwise in the "Notes" column.

Section A

| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | a |  | smooth curve passing through all error bars |  | 1 |
|  | b |  | $\begin{aligned} & x=2.5 \text { «cm» } \pm 0.2 \mathrm{~cm} \text { AND } \Delta x=0.5 \mathrm{~cm} \pm 0.1 \mathrm{~cm} \\ & \text { « } \frac{0.5}{2.5}=» 20 \% \checkmark \end{aligned}$ | Accept correctly calculated value from interval 15 \% to $25 \%$. | 2 |



| Question |  | Answers | Notes |  |
| :--- | :--- | :--- | :--- | :--- | :--- |
| $\mathbf{1}$ | $\mathbf{c}$ | ii | ALTERNATIVE 1 <br> $t^{-1}$ from $0.025 \mathrm{~s}^{-1}$ to $0.04 \mathrm{~s}^{-1} \checkmark$ <br> giving $t$ from 25 to $40 \checkmark$ | Do not allow ECF from MP1 <br> to MP2. |
| ALTERNATIVE 2 <br> the data do not support the hypothesis $\checkmark$ <br> any relevant support for the suggestion, eg straight line cannot be fitted through the error <br> bars and the origin $\checkmark$ | $\mathbf{2}$ |  |  |  |


| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 2 | a | i | refractive index $=1.5 \checkmark$ | Both correct value and 2SF required for [1]. | 1 |
|  | a | ii | $\begin{aligned} & \text { fractional uncertainty } x_{3}-x_{1}=\frac{0.04}{1.15}=0.035 \text { AND } x_{3}-x_{2}=\frac{0.04}{0.76}=0.053 \\ & \text { sum of fractional uncertainty }=0.088 \checkmark \\ & \text { «uncertainty }=\text { their } \mathrm{RI} \times 0.088 »=0.1 \checkmark \end{aligned}$ | Accept correct calculation using maximum and minimum values giving the same answer. | 3 |
|  | b | i | systematic error $\checkmark$ | Accept "zero error/offset". | 1 |
|  | b | ii | calculated refractive index is unchanged $\checkmark$ because both numerator and denominator are unchanged $\checkmark$ | Accept calculation of refractive index with 0.05 subtracted to each $x$ value. | 2 |
|  | c |  | numerator and denominator will be 10 times larger so refractive index is unchanged $\checkmark$ relative/absolute uncertainty will be smaller $\checkmark$ | "Constant material" is not enough for MP1. | 2 |

## Section B

| Option A - Relativity |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: |
| Question |  |  | Answers | Notes | Total |
| 3 | a |  | not being accelerated <br> OR <br> not subject to an unbalanced force <br> OR <br> where Newton's laws apply $\downarrow$ |  | 1 |
|  | b | i | $c \checkmark$ |  | 1 |
|  | b | ii | $c+v \checkmark$ |  | 1 |


| 4 | Y measures electrostatic repulsion only $\checkmark$ <br> protons are moving relative to $X$ «but not $Y$ » OR protons are stationary relative to $Y \checkmark$ <br> moving protons create magnetic fields around them according to $X \checkmark$ <br> $X$ also measures an attractive magnetic force OR relativistic/Lorentz effects also <br> present $\checkmark$ |  |  |
| :--- | :--- | :--- | :--- | :--- |


| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 5 | a |  | ALTERNATIVE 1 $\begin{aligned} & \text { «rest mass }=0.511 \mathrm{MeV} \mathrm{c}^{-2} » \gamma=\frac{2.30}{0.511}=4.50 \\ & v=c \sqrt{\frac{\gamma^{2}-1}{\gamma^{2}}} \text { OR } 3 \times 10^{8} \times\left(\frac{4.50^{2}-1}{4.50^{2}}\right)^{\frac{1}{2}} \checkmark \\ & 0.9750 c \checkmark \end{aligned}$ <br> ALTERNATIVE 2 $\begin{aligned} & \gamma=« \frac{1}{\sqrt{1-0.98^{2}}} \Rightarrow>5.0 \checkmark \\ & E=« \gamma m_{0} c^{2}=» 4.1 \times 10^{-13} \mathrm{~J} \checkmark \\ & E=2.6 \mathrm{MeV} \checkmark \end{aligned}$ | Must see answer to at least $3 S F$. | 3 |
|  | b | I | $\begin{aligned} & \text { distance }=\frac{0.800}{\gamma} \\ & 0.178 \mathrm{~m} \checkmark \end{aligned}$ | Accept 0.159 for $\gamma=5.0$. | 2 |


| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 5 | b | ii | $\begin{aligned} & \text { time }=\frac{0.800}{2.94 \times 10^{8}} \\ & 2.74 \text { ns } \checkmark \end{aligned}$ |  | 2 |
|  | b | iii | $\begin{aligned} & \frac{2.74}{4.5} \text { OR } \frac{0.178}{2.94 \times 10^{8}} \\ & 0.608 \mathrm{~ns} \checkmark \end{aligned}$ |  | 2 |
|  | b | iv | it is measured in the frame of reference in which both events occur at the same position OR <br> it is the shortest time interval possible $\checkmark$ |  | 1 |






| Question |  | Answers | Notes | Total |
| :--- | :--- | :--- | :--- | :--- |
| $\mathbf{7}$ | $\mathbf{a}$ | region of space with extreme/very large curvature of spacetime $\checkmark$ <br> such that light cannot escape the region $\boldsymbol{O R}$ escape speed within region is $>c \checkmark$ | Do not allow "large" or <br> omission of degree of <br> curvature. |  |
| $\mathbf{b}$ | time for 1 second spacecraft tick in observer frame $=1.07 \mathrm{~s} \checkmark$ <br> $1.07=\frac{1.00}{\sqrt{1-\frac{R_{\mathrm{S}}}{2.3 \times 10^{4}}}}$ OR $R_{\mathrm{S}}=2.96 \times 10^{3} \mathrm{~m} \checkmark$ <br> $M=« \frac{c^{2} \times 2.96 \times 10^{3}}{2 \times 6.67 \times 10^{-11}}=» 2.0 \times 10^{30} \mathrm{~kg} \checkmark$ |  |  |  |
| $\mathbf{2}$ |  |  |  |  |


| Option B - Engineering physics |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: |
| Question |  |  | Answers | Notes | Total |
| 8 | a |  | because $M g$ and $N$ act through the axis <br> OR <br> only $F$ has a non-zero lever arm «about the axis» |  | 1 |
|  | b | i | ALTERNATIVE 1 <br> use of Newton's law for linear motion: $M g \sin \theta-F=M a \checkmark$ use of Newton's law for rotational motion: $F R=I \alpha \checkmark$ combining $M g \sin \theta=M a+\frac{I \alpha}{R} \checkmark$ substitution of $I=\frac{1}{2} M R^{2}$ and $\alpha=\frac{a}{R} \checkmark$ to get result <br> ALTERNATIVE 2 $\begin{aligned} & M g h=\frac{1}{2} M v^{2}+\frac{1}{4} M v^{2} \text { «from } \frac{1}{2} I \omega^{2}=\frac{1}{2}\left(\frac{1}{2} M R^{2}\right) \frac{v^{2}}{R^{2}}> \\ & v^{2}=\frac{4}{3} g h \checkmark \\ & v^{2}=2 a s=2 a \frac{h}{\sin \theta} \checkmark \end{aligned}$ <br> manipulation to produce given answer $\checkmark$ | Accept correct use of torques about point of contact. | 4 |


| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 8 | b | ii | rearranging $s=\frac{1}{2} a t^{2}$ to get $t=\sqrt{\frac{2 s}{a}} \checkmark$ <br> substitution to get $t=$ « $\sqrt{\frac{2 \times 1.5}{\frac{2}{3} \times 9.81 \times \frac{1}{2}}} \gg 0.96 \mathrm{~s}$ |  | 2 |
|  | c |  | acceleration of ice is $g \sin \theta$ whereas for the solid cylinder acceleration is two thirds of this «so speed of ice must always be greater at same point» | Allow answers in terms of energies, eg ice does not use energy to rotate and therefore will have a greater translational speed. | 1 |
|  | d |  | the hollow cylinder has a greater moment of inertia $\checkmark$ and hence a smaller acceleration |  | 2 |


| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 9 | a | i | $1400 \mathrm{~K} \checkmark$ |  | 1 |
|  | a | ii | $\begin{aligned} & \frac{3}{2} P \Delta V=\frac{3}{2} \times 4 \times 10^{5} \times 3 \times 10^{-3} \\ & 1800 \mathrm{~J} \checkmark \end{aligned}$ |  | 2 |
|  | a | iii | $\begin{aligned} & 1800+P \Delta V=1800+4 \times 10^{5} \times 3 \times 10^{-3} \text { OR use of } \Delta Q=\frac{5}{2} P \Delta V \\ & 3000 \mathrm{~J} \checkmark \end{aligned}$ |  | 2 |
|  | a | iv | curve starting at A ending on line CB AND between B and zero pressure $\checkmark$ |  | 1 |
|  | b | i | $0 \checkmark$ |  | 1 |
|  | b | ii | ALTERNATIVE 1 <br> C has the same volume as B OR entropy is related to disorder $\checkmark$ higher temperature/pressure means greater disorder therefore entropy at C is greater «because entropy is related to disorder» <br> ALTERNATIVE 2 <br> to change from $B$ to $C, \Delta Q>0 \checkmark$ <br> so $\Delta S>0 \checkmark$ <br> $\Delta S$ related to disorder $\checkmark$ |  | 3 |


| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 10 | a |  | $« 118+105 \mathrm{kPa}\rangle=2.23 \times 10^{5} \mathrm{~Pa} \checkmark$ |  | 1 |
|  | b |  | ALTERNATIVE 1 <br> «from Bernoulli's Law» total pressure at $Q=$ static pressure + dynamic pressure $=$ constant « $2.2 \times 10^{5} \mathrm{~Pa}$ » $\checkmark$ <br> dynamic pressure $«=\frac{1}{2} \rho v^{2} »$ increases from zero, so static pressure decreases $\checkmark$ <br> ALTERNATIVE 2 <br> water rushes out of tap at higher velocity, so pressure is lower $\checkmark$ due to Bernoulli's Principle $\checkmark$ |  | 2 |
|  | c | i | $\begin{aligned} & R=\frac{1.27 \times 0.05 \times 1.00 \times 10^{3}}{1.8 \times 10^{-3}} \\ & R=3.5 \times 10^{4} \end{aligned}$ | Allow use of diameter to give $R=7.0 \times 10^{4}$ | 2 |
|  | c | ii | flow is turbulent $\checkmark$ | Answers in (c)(i) and (c)(ii) must be consistent. | 1 |


| Question |  | Answers | Total |  |  |
| :---: | :---: | :---: | :--- | :---: | :---: |
| $\mathbf{1 1}$ | $\mathbf{a}$ | high Q means low damping OR system oscillates with low damping $\checkmark$ <br> «exponential» decrease of amplitude/energy OR oscillates about 200 times before coming <br> to rest $\checkmark$ <br> loses about 3\% of energy per cycle OR loses small amount of energy each cycle $\checkmark$ | Notes |  |  |
|  | $\mathbf{b}$ | $\mathbf{i}$ | large amplitude/resonance $\checkmark$ |  |  |
|  | $\mathbf{b}$ | $\mathbf{i i}$ | small amplitude AND A «almost» in phase with P $\checkmark$ | $\mathbf{1}$ |  |


| Option C - Imaging |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: |
| Question |  | Answers | Notes | Total |
| 12 | a |  <br> one correct ray drawn <br> another correct ray $\downarrow$ <br> image located at intersection of rays, behind the mirror | Label I is required. | 3 |
|  | b | $\approx 0.4 \checkmark$ |  | 1 |
|  | c | image is in better focus/sharper $\boldsymbol{O R}$ parabolic do not suffer from spherical aberration $\checkmark$ parabolic mirrors reflect parallel rays through one point $\checkmark$ whereas spherical mirrors reflect parallel rays through different points $\checkmark$ | Award $3^{\text {rd }}$ mark even if implied in the answer. | 3 |


| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 13 | a |  | $\begin{aligned} & F_{\mathrm{o}}+f_{\mathrm{e}}=84 \text { so } f_{\mathrm{e}}=84-82=2 \mathrm{~cm} \\ & \text { «M }=\frac{f_{\mathrm{o}}}{f_{\mathrm{e}}}=\frac{82}{2}=» 41 \checkmark \end{aligned}$ |  | 2 |
|  | b |  | a sign convention is a way to distinguish between real and virtual objects or images or converging and diverging lenses $\checkmark$ |  | 1 |
|  | c | i | image will be virtual $v=-25 \mathrm{~cm} \checkmark$ $\begin{aligned} & \frac{1}{u}=\frac{1}{82}+\frac{1}{25} \checkmark \\ & «=19 \mathrm{~cm} \text { or } 0.19 \mathrm{~m} » \end{aligned}$ | Award [1 max] if $v=+25 \mathrm{~cm}$ used to give $u=-36 \mathrm{~cm}$. | 2 |
|  | c | ii | image will be real $v=84-19=65$ «cm» $\text { « } \frac{1}{u}=\frac{1}{2}-\frac{1}{65} » \text { so } u=2.1 \mathrm{~cm}$ |  | 2 |
|  | c | iii | $\begin{aligned} & M_{\mathrm{e}}=« \frac{D}{f_{\mathrm{e}}}+1=\frac{25}{82}+1 \Rightarrow 1.3 \text { AND } m_{\mathrm{o}}=« \frac{v}{f_{\mathrm{o}}}-1=\frac{65}{2}-1 \Rightarrow 31 \text { or } 32 \\ & \text { so } M=« M_{\mathrm{e}} m_{\mathrm{o}}=1.3 \times 31=» 40 \text { or } 41 \checkmark \end{aligned}$ | Far point adjustment gives $M=9.3$ (accept answers from interval 9.3 to 9.6), award [1 max] for full working. | 2 |


| Question |  | Answers | Notes | Total |
| :--- | :--- | :--- | :--- | :--- |
| $\mathbf{1 4}$ | a | curved, symmetrical path $\checkmark$ | Refraction on entry not <br> required and ignored in <br> diagram for simplicity. |  |
|  | b | waveguide dispersion means that rays not parallel to the central axis take longer to <br> transmit $\checkmark$ <br> in a graded-index fibre rays away from the central axis travel at a higher speed $\boldsymbol{O R}$ rays are <br> «refracted» closer to the central axis $\mathbf{O R}$ effective diameter of the fibre is reduced $\checkmark$ <br> because refractive index is greater in the centre $\mathbf{O R}$ refractive index is less at the edge $\checkmark$ |  |  |


| 15 | a | i | $\mu=2.7 \times 10^{-3}\left( \pm 0.3 \times 10^{-3}\right)^{\checkmark}$ <br> So $\left.\frac{I}{I_{0}}=« e^{-\mu x}=e^{-\left(2.7 \times 10^{-3} \times 8 \times 10^{-2}\right)}\right\rangle=0.9999 \approx 1.0 \checkmark$ | 2 |
| :---: | :---: | :---: | :---: | :---: |
|  | a | ii | « $\mu=50$ to give $\frac{I}{I_{0}}=1.8 \times 10^{-2} \checkmark$ | 1 |
|  | b |  | low energy radiation removed but not high energy radiation radiation has narrower range of energies $\checkmark$ only necessary radiation reaches the patient making it safer $\checkmark$ | 2 max |


| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 16 | a |  | Advantage: no ionizing radiation OR high res images of soft tissue OR 3D images $\checkmark$ <br> Disadvantage: «generally" more expensive $\mathbf{O R}$ takes much longer $\boldsymbol{O R}$ less detail of bony structures «than X-ray» OR noisy for patient OR claustrophobic for patient OR cannot be used for patients with metal implants $\checkmark$ | Do not accept advantages that are also true of $X$-rays, eg non-invasive. | 2 |
|  | b |  | a gradient field is added to a magnetic field that is originally uniform across patient $\checkmark$ the gradient field varies linearly across patient as the protons relax a «radio frequency» signal is emitted $\checkmark$ the emitted signal frequency is proportional to the total strength of the magnetic field $\checkmark$ the signal frequency depends on the emission position in the patient |  | 3 max |



| 18 | a | $\begin{aligned} & T=\frac{2.9 \times 10^{-3}}{740 \times 10^{-9}} \\ & 3900 \mathrm{~K} \checkmark \end{aligned}$ | Answer must be to at least 2SF. | 2 |
| :---: | :---: | :---: | :---: | :---: |
|  | b | $\begin{aligned} & L=5.67 \times 10^{-8} \times 4 \pi \times\left(3.1 \times 10^{10}\right)^{2} \times 4000^{4} \\ & =1.8 \times 10^{29} \mathrm{~W} \end{aligned}$ | Accept use of $3900^{4}$ to give $1.6 \times 10^{29} \mathrm{~W}$ | 2 |
|  | c | absorption lines in spectra $\checkmark$ are specific to particular elements | Accept "emission lines in spectra". | 2 |
|  | d | helium $\checkmark$ |  | 1 |


| Question |  | Answers |  | Notal |
| :--- | :--- | :--- | :--- | :--- | :--- |
| $\mathbf{1 8}$ | $\mathbf{e}$ | helium flash $\checkmark$ <br> expansion of outer shell OR surface temperature increase $\checkmark$ <br> planetary nebula phase $\checkmark$ <br> only the core remains $\checkmark$ <br> if below $1.4 \mathrm{Ms} /$ Chandrasekhar limit then white dwarf $\checkmark$ |  |  |


| 19 | a | i | $z=\frac{\Delta \lambda}{\lambda_{0}}$ where $\Delta \lambda$ is the redshift of a wavelength and $\lambda_{0}$ is the wavelength measured at rest on Earth $\boldsymbol{O R}$ it is a measure of cosmological redshift $\checkmark$ | Do not allow just "redshift". | 1 |
| :---: | :---: | :---: | :---: | :---: | :---: |
|  | a | ii | «z $=\frac{R}{R_{\circ}}-1, \frac{R_{0}}{R}=\frac{1}{z+1}$ 》 so $\frac{R_{0}}{R}=« \frac{1}{1.16} 》=0.86 \checkmark$ | Do not accept answer 1.16. | 1 |
|  | a | iii | $\begin{aligned} & v=z c=0.16 \times 3 \times 10^{8}=4.8 \times 10^{4} « \mathrm{~km} \mathrm{~s}^{-1} » \checkmark \\ & d=\frac{v}{H_{0}}=\frac{4.8 \times 10^{4}}{68}=706 \mathrm{Mpc} \text { OR } 2.2 \times 10^{25} \mathrm{~m} \end{aligned}$ |  | 2 |
|  | b |  | as the universe expanded it cooled/wavelength increased $\checkmark$ <br> the temperature dropped to the present approximate $3 \mathrm{~K} \mathbf{O R}$ wavelength stretched to the present approximate $1 \mathrm{~mm} \checkmark$ | Value is required for MP2. | 2 |


| Question |  | Answers | Total |
| :--- | :--- | :--- | :--- | :--- | :--- |
| $\mathbf{2 0}$ | a | a gas cloud will collapse to form a star $\checkmark$ <br> if «the magnitude of» the gravitational potential energy of the particles is greater than the <br> kinetic energy of the particles OR mass of the cloud is greater than the Jeans mass $\checkmark$ | Notes |
|  | $\mathbf{b}$ | la have consistent maxima in their light curves but II vary $\checkmark$ <br> la has a strong ionized Sill line but II has hydrogen lines in their spectra $\checkmark$ <br> la was a white dwarf but II are massive stars $\checkmark$ <br> la form from binary systems but II are the result of core collapse of a star $\checkmark$ <br> la can be used as standard candles but II are not $\checkmark$ | $\mathbf{3}$ |


| Question |  |  | Answers | Notes | Total |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 21 | a | i | curve beginning on $R=0$ before present time and ending after present time on $R=0 \checkmark$ |  | 1 |
|  | a | ii | curve starting earlier than $C$ with general shape shown above $\checkmark$ coincides with curve $C$ at present time $\checkmark$ | Judge by eye. | 2 |


| Question |  | Answers | Notes |
| :--- | :--- | :--- | :--- | :--- | :--- |
| $\mathbf{2 1}$ | $\mathbf{b}$ | rotation speeds of galaxies is greater at the edges than expected $\checkmark$ <br> so the density at the edges must be greater than that supplied by luminous matter <br> alone $\checkmark$ | Accept any other valid piece <br> of evidence, eg gravitational <br> lensing, which provides a <br> good measure of galactic <br> cluster masses. |

